

Checking Out What is Checked In

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- Need to optimize balance of checking between users and developers.
 - infuse more of a culture of detailed checking by developers.
 - add help through analysis group. Today: review list, brainstorm additional items, and gather volunteers.
- Start a more formal list of some of the things to check. Appended list is not complete!

First Principles

- When a developer checks in a change, the new module should be tested *as it will be used.*
 - compiling, linking, and completing 100 events is certainly not bad...but it's not enough!
 - find the memory leaks on a production-scale job and confirm that PC users can still run large jobs without 10GB of RAM. [Related: what is the status of the “recent” RootIO leak?]
 - check the stability and functionality both in generation and event read-back. check that the single event display works.
 - think through the impacts, and check the system test histograms...and any other distributions that might be affected... before/after change. If help is needed, ask for it!
- Follow your nose. The suggested list is just a minimal place to start. If you notice anything odd, grab on and don't let go until you have an answer. Then, share it.

Some Things to Check (I)

- Event Display. Scan a few hundred events:
 - any laws of physics obviously violated (energy, momentum, charge conservation)?
 - TKR hits match charged particle trajectories in detail [magnify regions!]. recon tracks reasonably match MC truth.
 - CAL recon locations sensible?
 - particle flux is as expected (direction, energy, type)?
 - compare trigger bits with display in detail.
 - ACD hits consistent with MC truth particle trajectories?

Some Things to Check (II)

■ Trigger Distributions

- rates of TKR, CAL-LO and CAL-HI for benchmark fluxes (suggest `all_gamma`, `backgndavg`, and `normal_gamma_10GeV`)
- all 32 filter status bit frequencies (now there, should be declared stable very soon)
- A_{eff} at trigger level (require L1T) at 100 MeV, 1 GeV, 10 GeV at \sim normal incidence and at $\sim 50^\circ$ (better: plot A_{eff} vs θ).
- After L1T, total visible raw energy in CAL for benchmark gamma fluxes and total background flux.

Some Things to Check (III)

- Basic recon distributions for reference fluxes and for a run of “empty” events (checks noise implementations), requiring a L1T:
 - #hit ACD tiles, and frequency of hits for each tile
 - #TKR hits by layer (and by tower?)
 - #CAL logs hit. Total raw CAL energy visible.
- When the “standard” analysis is ready, check final PSF and A_{eff} + FOV for reference fluxes. Look at A_{eff} by layer. Residual background rate, in Hz, by flux component. Total residual background raw visible energy in CAL, and reconstructed energy. Check diffuse_gamma rate after all cuts, and plot reconstructed energy.