



Recall: DC1 December Workshop

Something very significant started....



post hoc ergo propter hoc?



Purposes of the Data Challenges

- "End-to-end" testing of analysis software.
- Familiarize team with data content, formats, tools and realistic details of analysis issues (both instrumental and astrophysical).
- If needed, develop additional methods for analyzing LAT data, encouraging alternatives that fit within the existing framework.
- Provide feedback to the SAS group on what works and what is missing from the data formats and tools.
- Uncover systematic effects in reconstruction and analysis.

Support readiness by launch time to do all first-year science.



Data Challenge Planning Approach

- Walk before running: design a progression of studies.
- DC1. Modest goals. Contains most essential features of a data challenge.
 - 1 simulated day all-sky survey simulation
 - find the sources, including GRBs
 - a few physics surprises
 - exercise:
 - exposure, orbit/attitude handling, data processing pipeline components, analysis tools
- DC2, start end of CY04. More ambitious goals. ~One simulated month.
 - based on DC1 lessons
 - discussion tomorrow
- DC3. Support for flight science production.





Eye on the ball

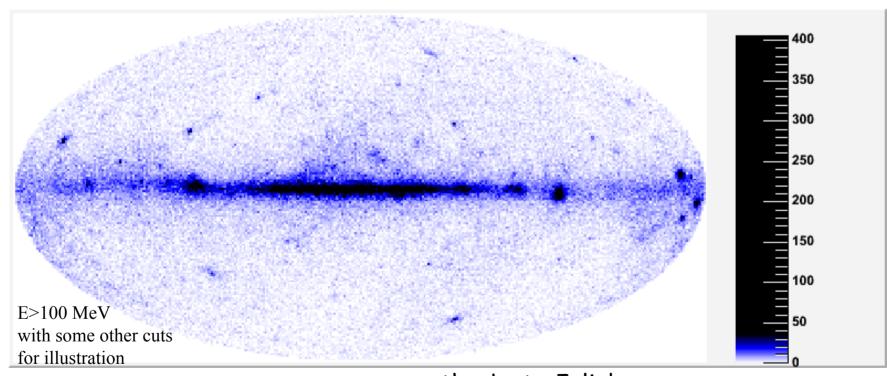
This would be very bad...





The DC1 Sky

One day all-sky survey. Generated E>20 MeV.



thanks to Julie!

Lots to analyze!
A few surprises to find...

Now the tools are in place to do individual studies after DC1!



DC1 Minimum Results

- The existence of the data sets and the volume of data generated for background analyses already meets one of the success criteria.
- All reports of experiences welcome and encouraged
- A minimum set of plots and tables that we must collectively produce:
 - TABLE 1: found sources, ranked by flux (E>100 MeV). Table has the following columns
 - reconstructed location and error circle
 - flux (E>100 MeV) and error
 - significance
 - 3EG identification (yes or no) [note: DON'T assume DC1 sky is the 3EG catalog!]
 - extra credit:
 - » include flux below 100 MeV
 - » spectral indices of brightest sources
 - » comparison of 3EG position and flux characteristics with GLAST analysis
 - FIGURE 1: LogN-logs plot of TABLE1
 - TABLE 2: list of transients detected. Columns are
 - location and error circle
 - flux (E>100 MeV) and error
 - significance
 - duration
 - FIGURE 2: light curve
 - Extra credit: FIGURE 2a: spectra.
- S. Ritz PLUS: reports of any physics surprises found.



Beyond the minimum

- Here were a few suggestions:
 - generated and released more days of data
 - better exercise tools and infrastructure
 - more transients
 - spectral analyses
 - localization studies
 - one-day localization of Vela is particularly interesting
 - analysis improvements
- But don't let this list limit you. The sky is the limit!
- Studies should continue post DC1. Wiki will persist.



Summary

- DC1 is already a great success, due to the hard work of many people working together. We have learned many lessons already.
- Now the fun part: a parade of results!
- Note: we want to hear from everyone who participated in the data challenge.
 - learning from the efforts that did not result in talks today is a key goal of DC1!!